

Equations of Motion of the Asteroid Moving in Earth's Atmosphere

$$\frac{d\vec{v}}{dt}(\vec{r}) = \vec{g} - \frac{c_d}{2} \frac{\rho_{at}(x) \vec{v}(\vec{r})^2 \pi R(t)^2}{m} \vec{e}_t$$

Equation of Deformation of the Ellipsoid Shaped Asteroid

$$\frac{dv_q}{dt} = \frac{\rho_{at}(x)}{\rho_{imp}} \frac{\vec{v}(\vec{r})^2}{R(t)}$$

Density of the Exponential Atmosphere (isotherm)

$$\rho_{at}(x) = \rho_0 \exp\left(-\frac{m_p g x}{k_b T_{amb}}\right)$$

Conservation Equations for a Gas Dynamic Flow Field

$$\begin{aligned} \frac{\partial}{\partial t} \rho + \vec{\nabla} \cdot \rho \vec{v} &= 0 \\ \frac{\partial}{\partial t} \rho \vec{v} + \vec{\nabla} \cdot \rho \vec{v} \otimes \vec{v} &= -\vec{\nabla} p + \rho \vec{g} \\ \frac{\partial}{\partial t} e + \vec{\nabla} \cdot (e + p) \vec{v} &= \rho \vec{g} \cdot \vec{v} \\ p &= (\kappa - 1) \left(e - \frac{1}{2} \rho \vec{v}^2 \right) \end{aligned}$$

Conservation Equations for the Hypervelocity Impact

$$\begin{aligned}
 \frac{\partial}{\partial t} \rho_i + \vec{\nabla} \cdot \rho_i \vec{v} &= 0, \quad i \in \{1, 2, 3\} \\
 \rho &= \rho_1 + \rho_2 + \rho_3 \\
 p &= p_1 + p_2 + p_3 \\
 \frac{\partial}{\partial t} \rho \vec{v} + \vec{\nabla} \cdot \rho \vec{v} \otimes \vec{v} &= -\vec{\nabla} p + \rho \vec{g} \\
 \frac{\partial}{\partial t} e + \vec{\nabla} \cdot (e + p) \vec{v} &= \rho \vec{g} \cdot \vec{v} \\
 \varepsilon_i &= \frac{\rho_i}{\rho} e - e_i, \quad i \in \{1, 2\} \\
 \frac{\partial}{\partial t} \varepsilon_i + \vec{\nabla} \cdot \varepsilon_i \vec{v} &= \left(p_i - \frac{\rho_i}{\rho} p \right) \vec{\nabla} \cdot \vec{v}, \quad i \in \{1, 2\} \\
 e_3 &= e - (e_1 + e_2)
 \end{aligned}$$

Governing Equations for an Incompressible Flow Field for Large Density Contrasts

$$\begin{aligned}
 \frac{\partial}{\partial t} \rho + \vec{\nabla} \cdot \rho \vec{v} &= 0 \\
 \frac{\partial}{\partial t} \rho + (\vec{\nabla} \rho) \cdot \vec{v} &= 0 \\
 \vec{\nabla} \cdot \vec{v} &= 0 \\
 \frac{\partial}{\partial t} \vec{v} + \vec{v} \cdot \vec{\nabla} \vec{v} &= -\frac{1}{\rho} \vec{\nabla} p + \vec{g} \\
 \vec{\nabla} \cdot \left(\frac{1}{\rho} \vec{\nabla} p + \vec{v} \cdot \vec{\nabla} \vec{v} \right) &= 0 \\
 \vec{\nabla} \cdot (\vec{\nabla} \Omega - \vec{v}) &= 0 \\
 p &= p + \frac{d\Omega}{dt} \\
 \vec{v} &= \vec{v} - \vec{\nabla} \Omega
 \end{aligned}$$

Shallow Water Equations

$$\begin{aligned}\frac{\partial h}{\partial t} + \vec{\nabla} \cdot (h \vec{v}) &= 0 \\ \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} + g \vec{\nabla} (h + s) &= 0\end{aligned}$$

One Dimensional Shallow Water Equations on a Sphere

$$\begin{aligned}\frac{\partial h}{\partial t} + \frac{1}{R_E} \left(\frac{\partial}{\partial \theta} (h v) + \cot(\theta) h v \right) &= 0 \\ \frac{\partial v}{\partial t} + \frac{1}{R_E} \left(v \frac{\partial v}{\partial \theta} + g \frac{\partial h}{\partial \theta} \right) &= 0\end{aligned}$$

Displacement Energy and Volume of Parabolic Water Cavity

$$\begin{aligned}E_{cv} &= \frac{\pi}{6} \rho_w g R^2 t^2 \\ V_{cv} &= -\frac{\pi}{2} t R^2\end{aligned}$$

Energy and Volume of Breakwater in Worse Case Estimation

$$\begin{aligned}E_{ts} &= \frac{\pi}{3} \rho_w g R_E \sin \left(\frac{b}{R_E} \right) (5 h^3 + d^3 \tan^2(\alpha)) \\ V_{ts} &= \pi R_E \sin \left(\frac{b}{R_E} \right) (2 h^2 - d^2 \tan(\alpha))\end{aligned}$$

Nomenclature

θ	latitude	\vec{e}_t	tangential unit vector
E	energy	e	energy density
\vec{g}	gravity acceleration	h	water height, breakwater height
k_b	Boltzmann constant	κ	adiabatic coefficient
p	pressure	R	radius
\vec{r}	spatial vector	ρ	density
\vec{v}	velocity	x	atmospheric height
d	low tide distance	s	ocean bottom topography
m	mass	T	temperature
c	coefficient	Ω	gauge potential
t	depht	V	volume
b	distance	\vec{r}	spatial vector

Subscripts

amb	ambient	at	atmospheric
d	drag	E	earth
imp	impactor	p	atmospheric particle
ts	tsunami	0	ground level
w	water	t	tangent
cv	cavity	q	ellipsoid boundary